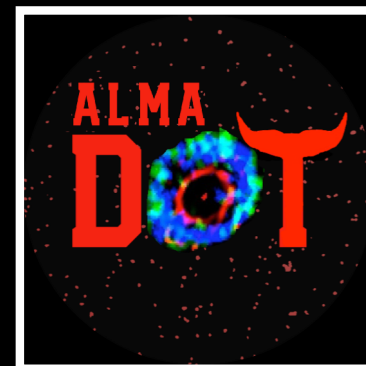




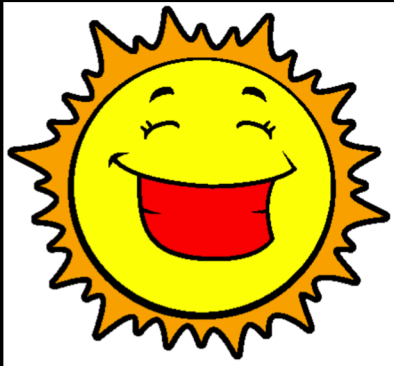
AstroChemical Origins (ACO) H2020 EU Innovative Training Network

*Claudio Codella
INAF – Osservatorio Astrofisico di Arcetri*



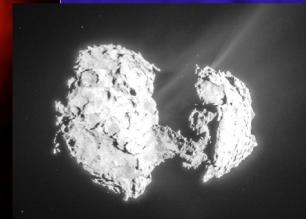
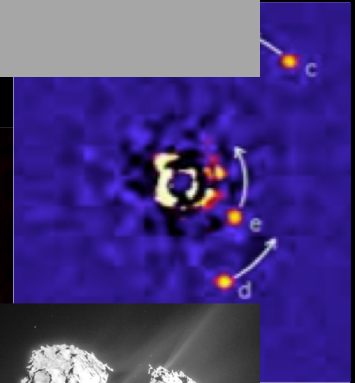
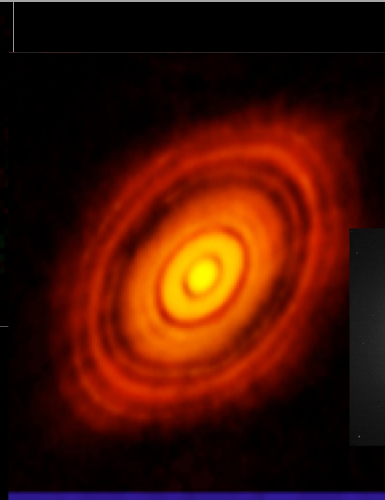
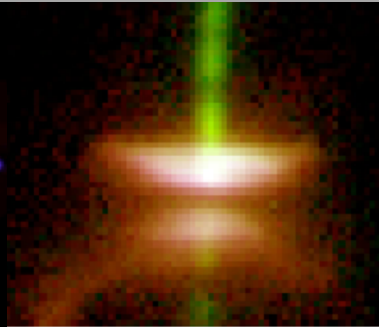
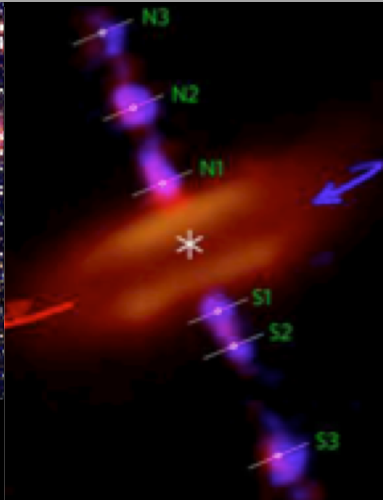
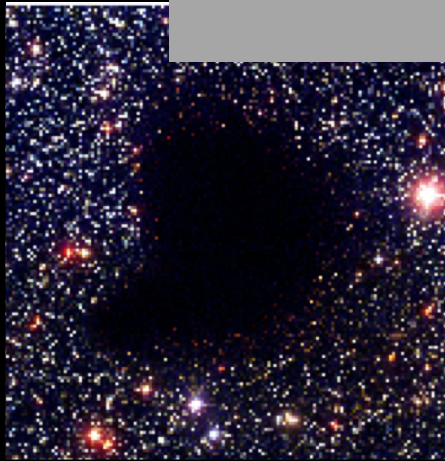


THE ROOTS OF ACO

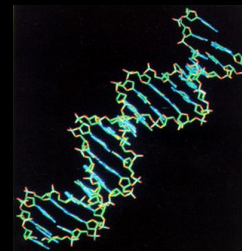
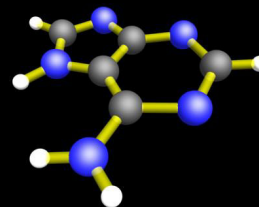
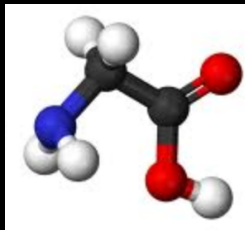
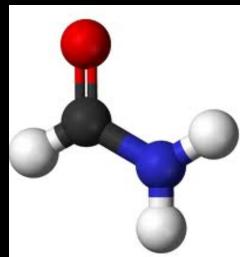
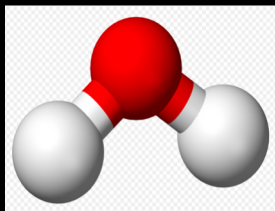


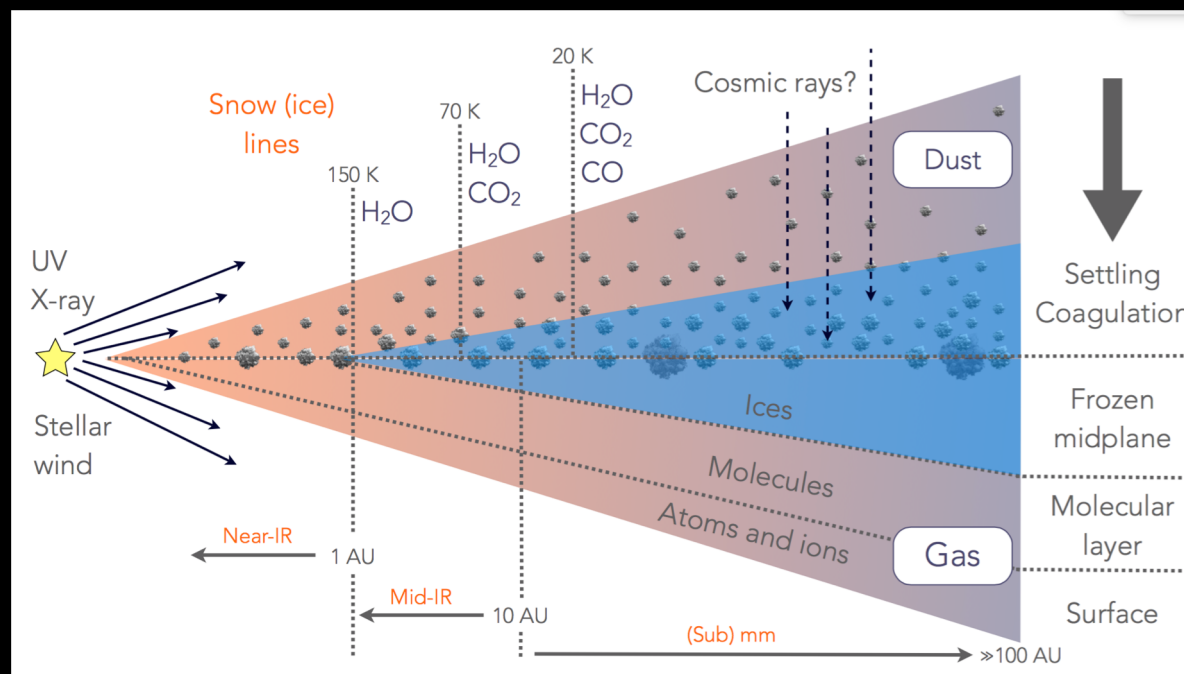
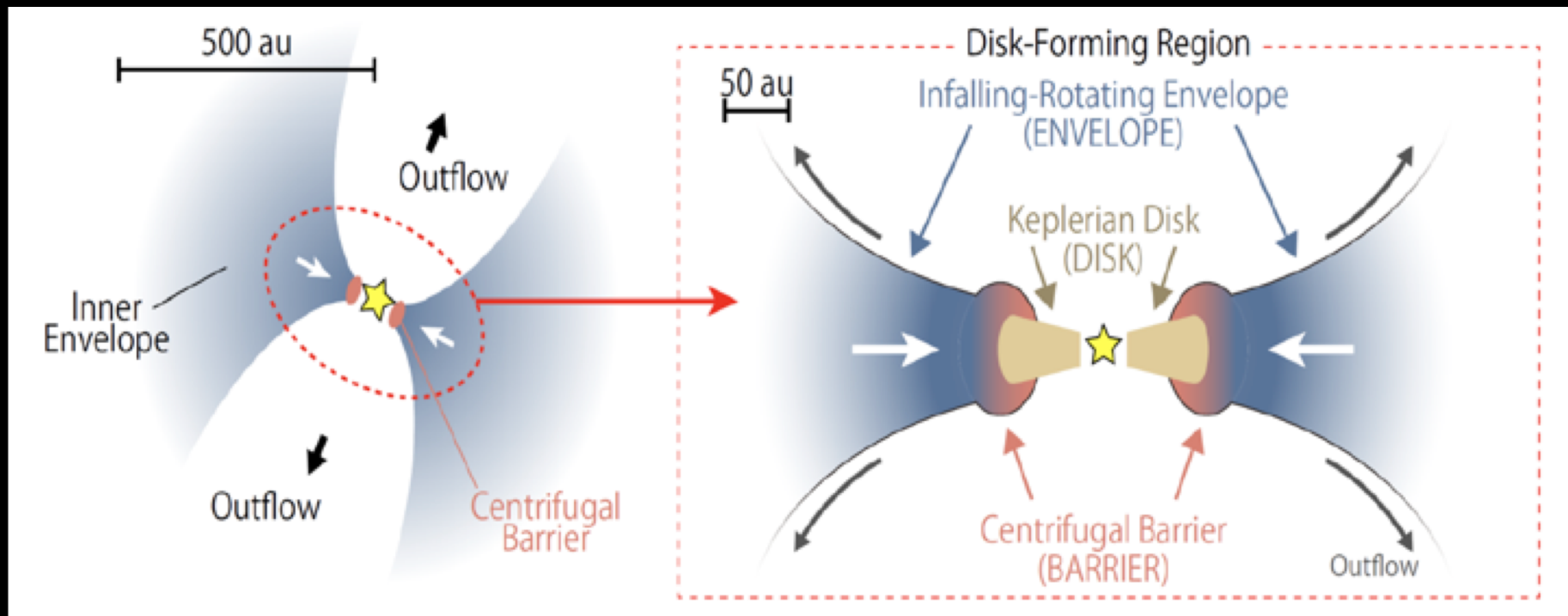
*The formation of a Sun-like star,
and its Solar System.
The emergency of life*

Planetary composition: disk chemical reset
or inheritance ?

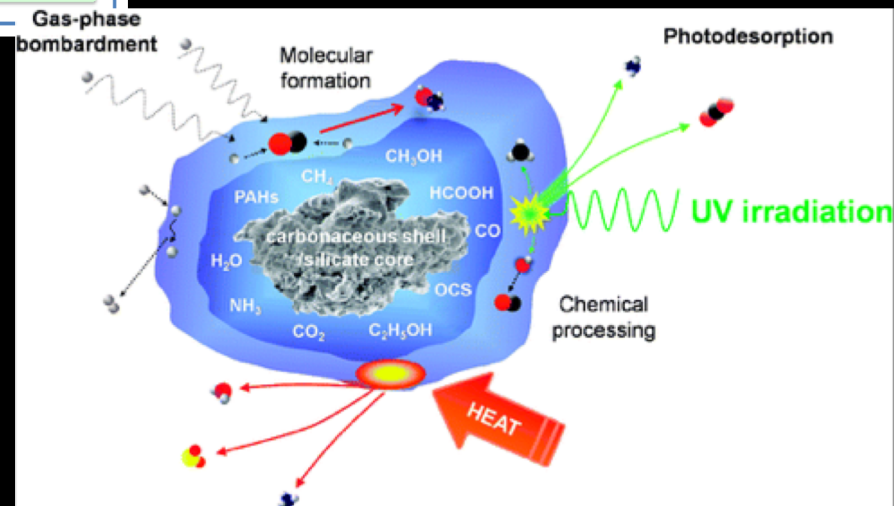
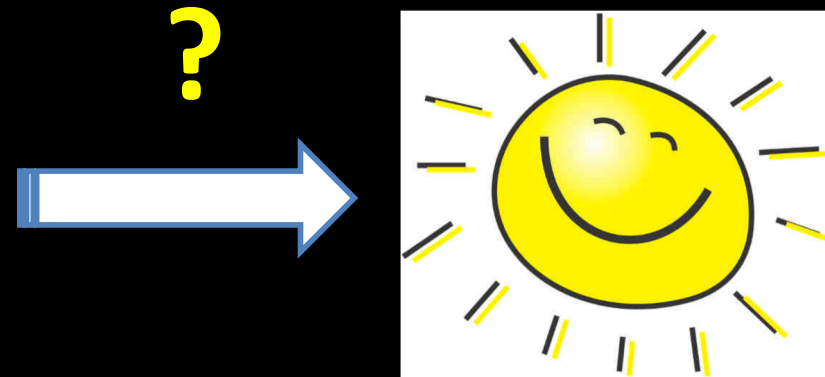
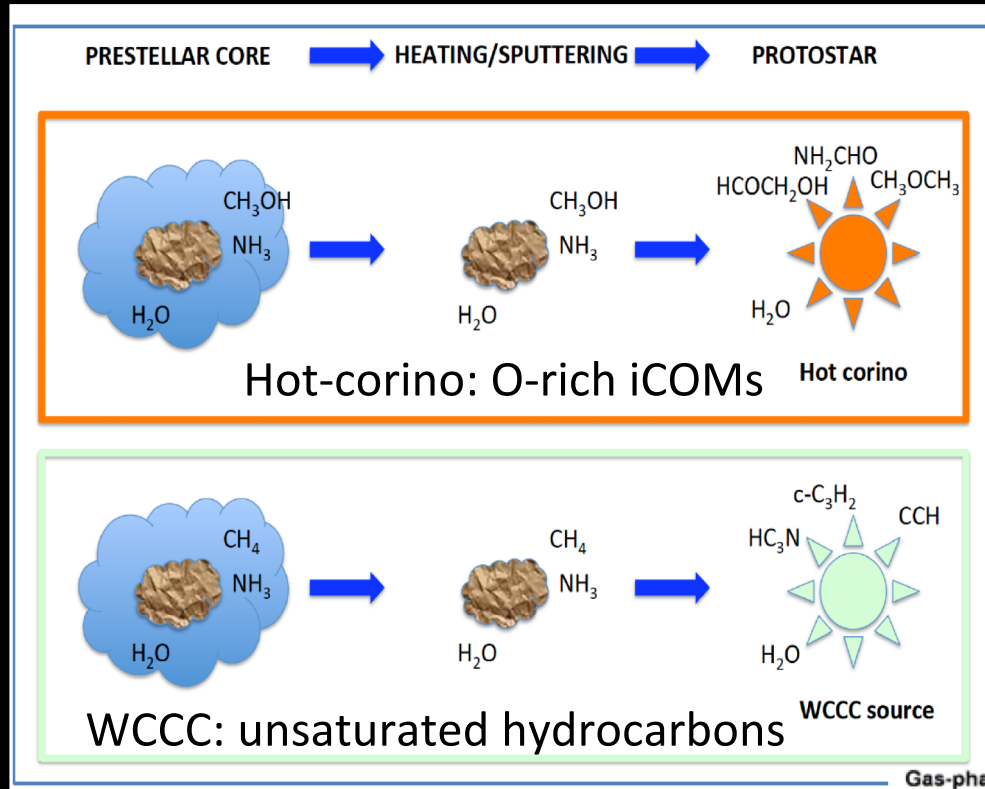


Time

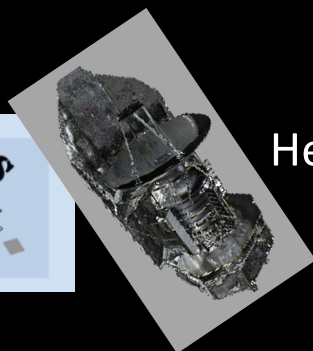




Tracing ice mantle history in Solar-type protostars



Tracing our chemical origins: Astrochemistry of forming Sun-like star



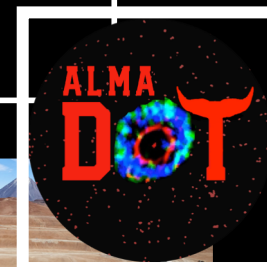
Herschel



IRAM 30-m



- 1 Premiale INAF
- 6 Large Programs
- 2 Astrofit
- 1 SIR project
- 1 PRIN-MIUR
- 1 PRIN-INAF
- 1 H2020 EU-ITN



ALMA

Today

Time



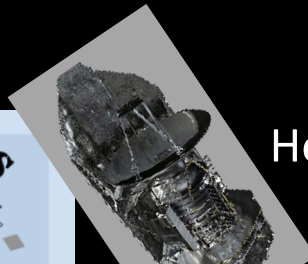
IRAM PdBI



IRAM NOEMA



Tracing our chemical origins: Astrochemistry of forming Sun-like star



Herschel

1 Premiale INAF
6 Large Programs
2 Astrofit
1 SIR project
1 PRIN-MIUR
1 PRIN-INAF
1 H2020 EU-ITN

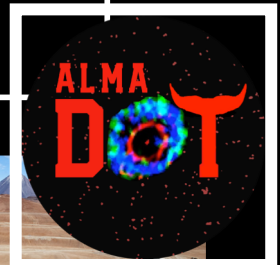
2009

2 Researchers TI @INAF
L. Podio, D. Fedele

6 PhDs @INAF
E. Bianchi, M. De Simone,
M. Corazzi, R.G. Urso,
S. Mercimek, L. Evans



IRAM PdBI



IRAM NOEMA



ALMA

Today



GENESIS - SKA PI C. Codella FOUR PILLARS



1. PLANET FORMATION:
Models, simulations, & observations

2. VOLATILES EVOLUTION:
Complex Organics as the building
blocks of life
(quantum-mechanical computations
of gas reactions)

3. LABORATORY EXPERIMENTS

4. COMMUNICATION AND
DISSEMINATION

www.genesis.inaf.it
(2018-2021: 98 papers...)



Four laboratories involved in the study of the formation and survival of complex molecules in space

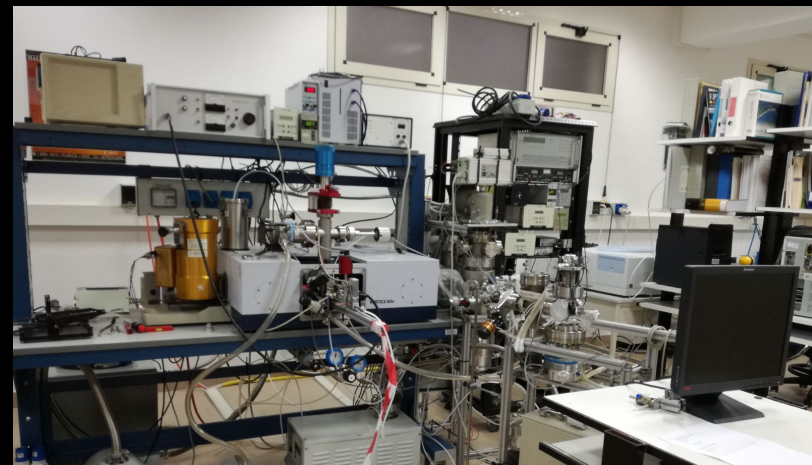
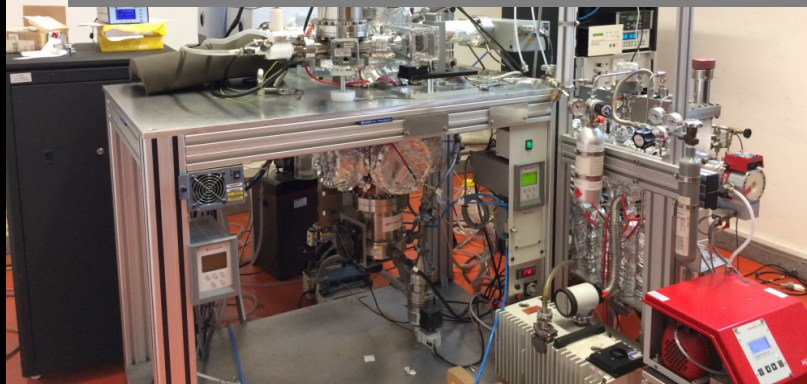


Laboratorio di Astrofisica
Sperimentale di Catania
M.E. Palumbo

Laboratorio di Astrobiologia Arcetri
J.R. Brucato



LIFE Light Irradiation Facility for
Exochemistry, Palermo
A. Ciaravella



Laboratorio di Fisica Cosmica Capodimonte
V. Mennella



ACO

ACO team
Total: 4MEuro
INAF: 400 kEuro



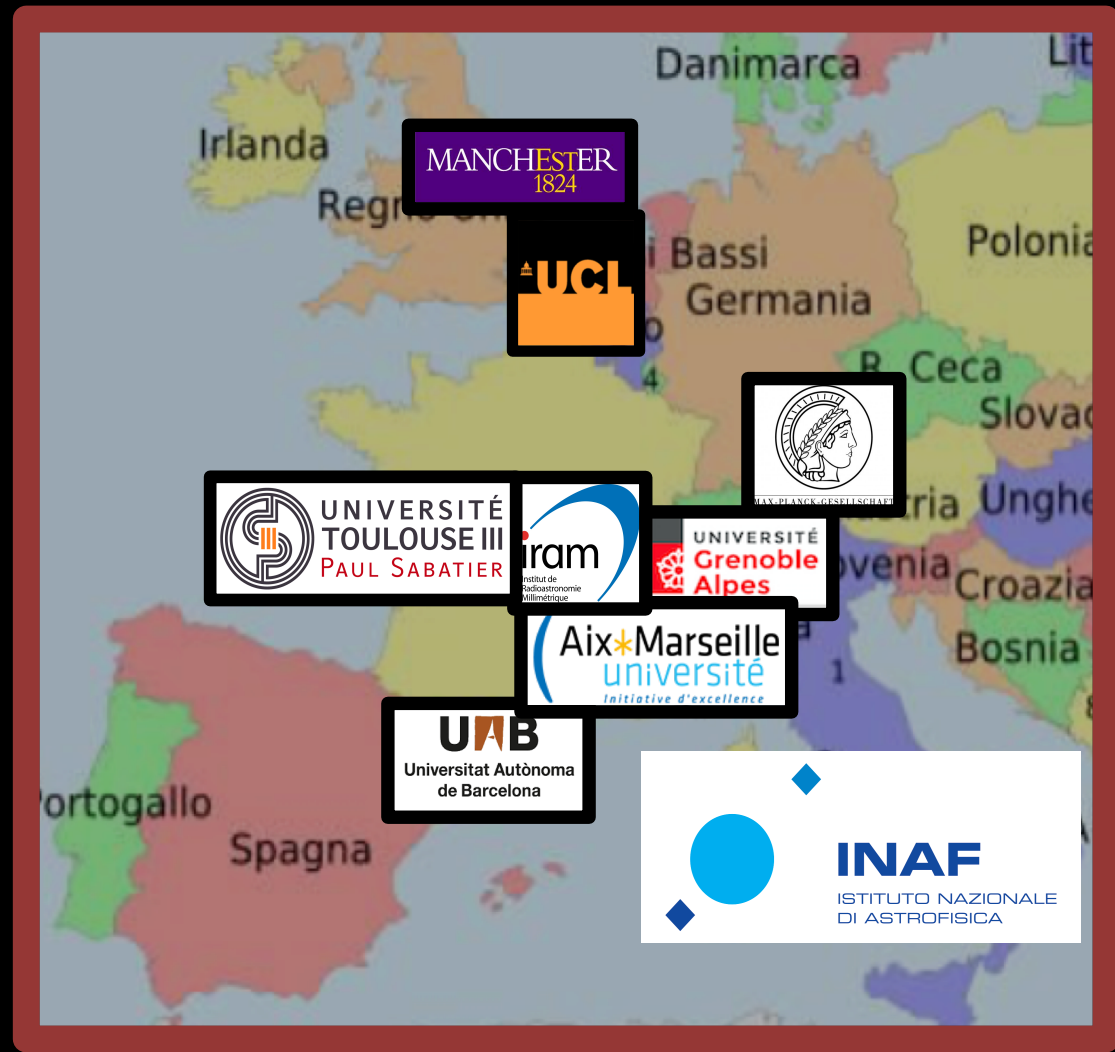
BENEFICIARIES	COORDINATOR
Un. Grenoble Alpes	PI: Ceccarelli
Un. Sabatier Toulouse	Vastel
Un. Aix-Marseille	Theulé
Istituto Nazionale Astrofisica	Codella
Un. Torino	Ugliengo
Un. Perugia	Balucani
Un. Trento	Ascenzi
Un. College London	Viti
Un. Manchester	Piccirillo
Un. Auton. Barcelona	Rimola
IRAM	Neri
POAM Electronics	Amiri
MASTER-UP	Faginas-Lago

PARTNER ORGANIZATION

Vastalla Srl	Software
MPE Garching	
Mellanox Tech.	High Performance Computing
Springer Nature	Publishing Company
PathControl	Oil drilling: Matlab software
RobertHunter	Website Designers
Oritur	Risk Management Consultant
Aethia Srl	High Performance Computing

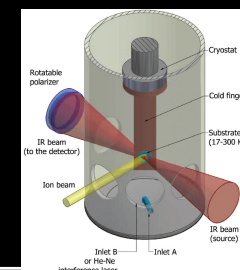
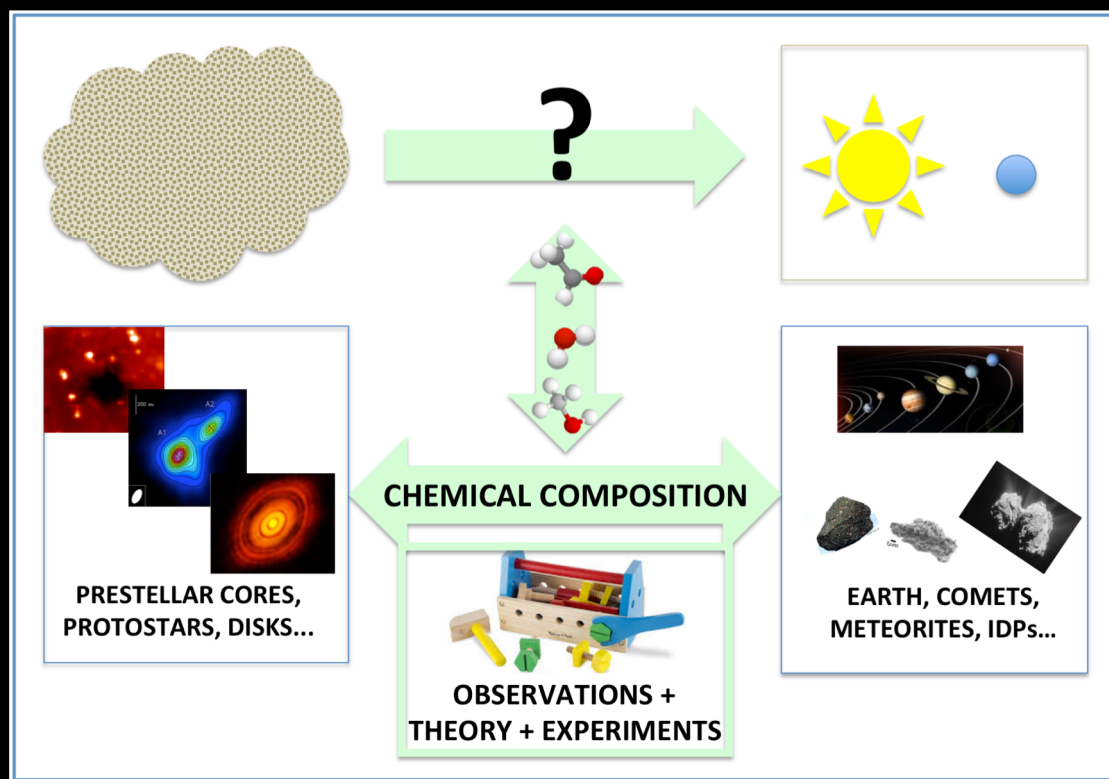
Electronics/Receivers
Computations/Chemistry

Academic
Non-academic

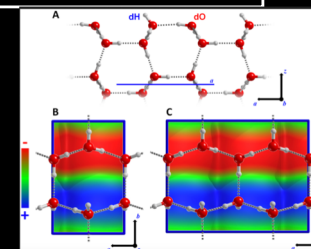
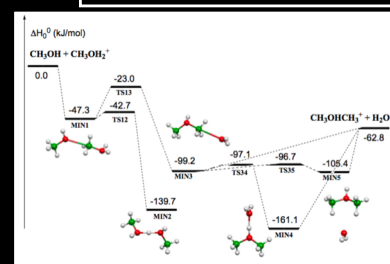


Scientific goals

USE THE CHEMICAL COMPOSITION AS A TOOL TO RECONSTRUCT THE EARLY PHASES OF THE SOLAR SYSTEM FORMATION



- (i) laboratory outputs,
- (ii) surface simulations,
- (iii) quantum-mechanical computations

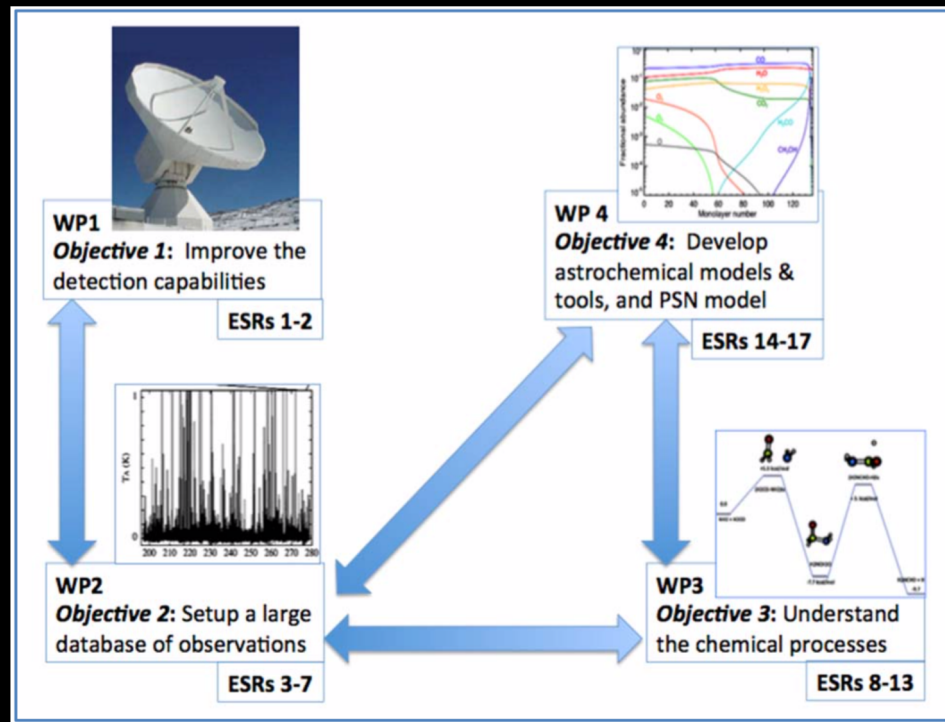


→ OBSERVE THE CHEMICAL PROXIES LINKED TO THE SOLAR SYSTEM PRIMITIVE OBJECTS TOWARDS PRESTELLAR CORES, PROTOSTARS AND PROTOPLANETARY DISKS

→ AND INTERPRET THEM VIA ASTROCHEMICAL MODELS

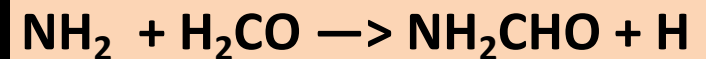
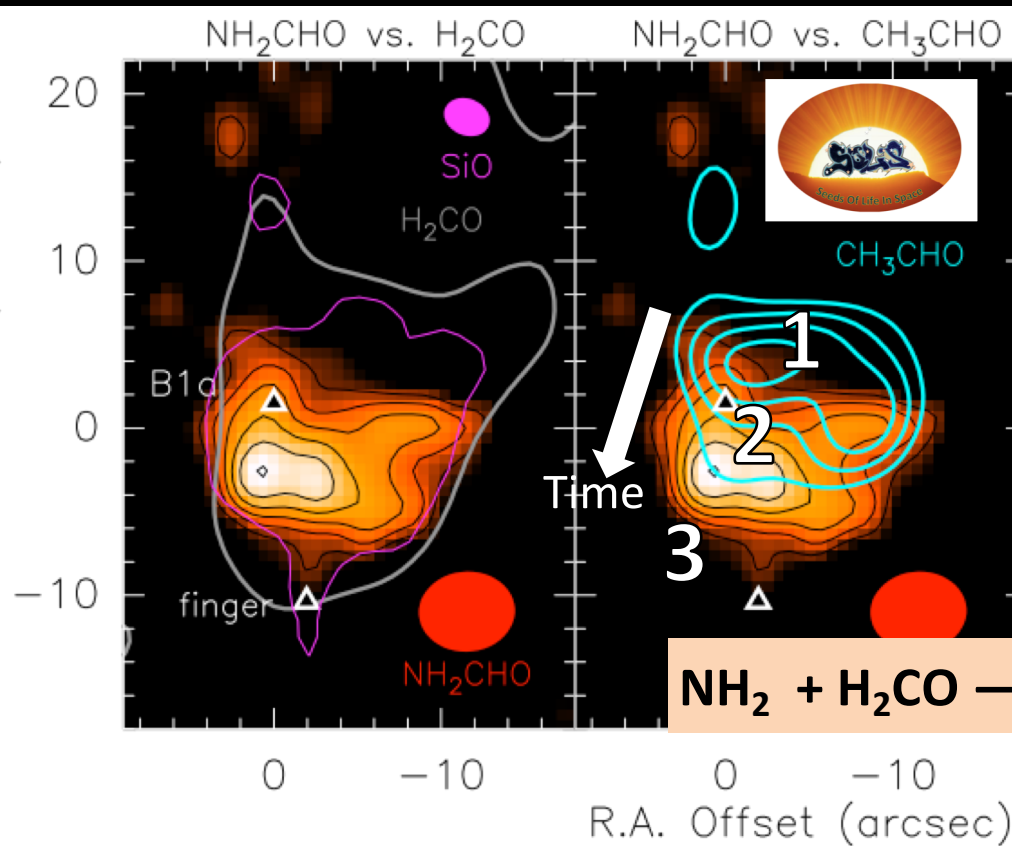
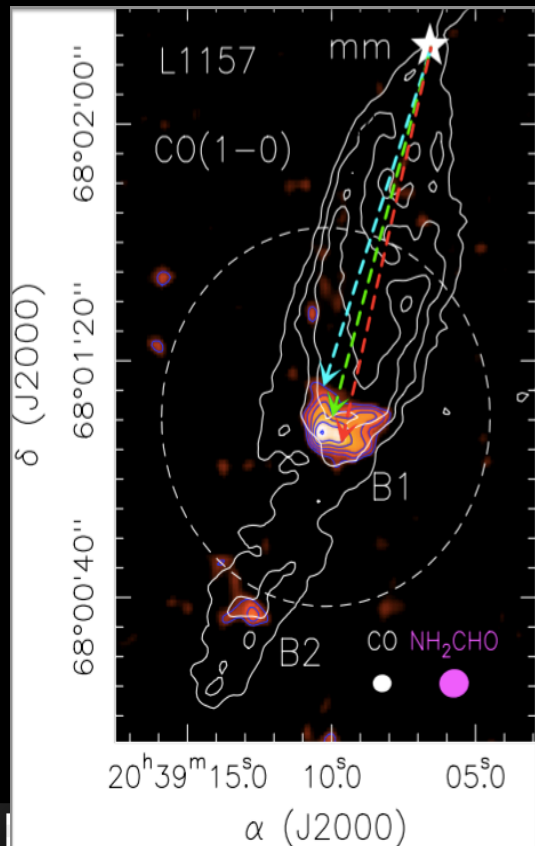
ACO FOUR SPECIFIC OBJECTIVES

*THE ACO SCIENTIFIC PROJECT WILL BE CARRIED OUT
BY 17 EARLY STAGE RESEARCHERS (ESRs)*



- 3 NETWORK SCHOOLS
- IMMERSIVE VIRTUAL REALITY SCHOOL
- 2 CONFERENCES
- standard courses at host universities

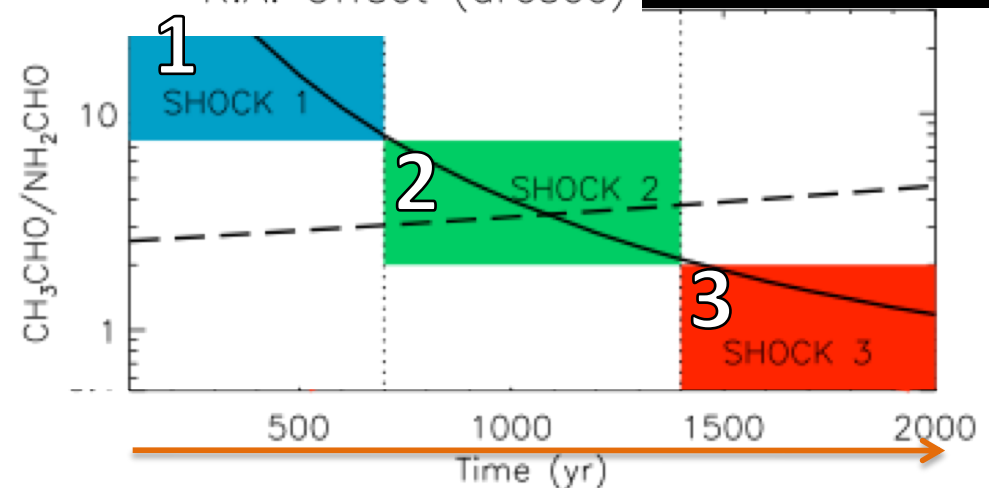
Astro + Chemictry = Astrochemistry !



Codella et al. (2017)

The smoking gun
of gas-phase at work
(at least in L1157-B1)

Barone et al. (2015)
Vazart et al. (2016)
Skouteris et al. (2017)





The first ALMA LP on Astrochemistry

The FAUST
(Fifty AU Study of Protosun Analogues)
ALMA Large Program



C. Ceccarelli (IPAG, Grenoble)
C. Codella (INAF-Arcetri)



S. Yamamoto (Tokyo University)
N. Sakai (RIKEN)

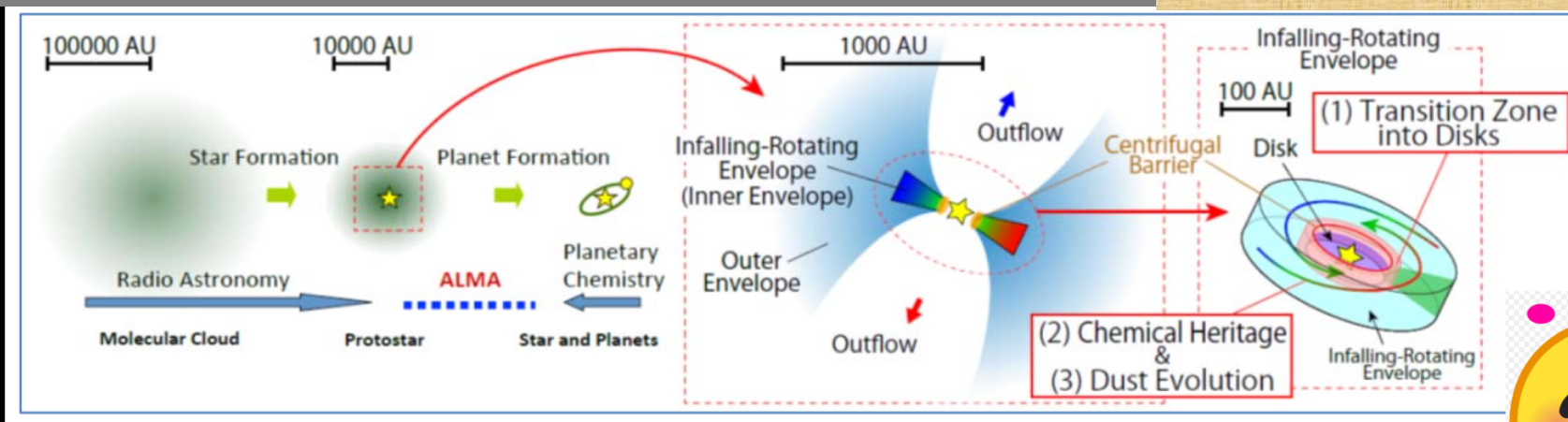


C. Chandler (NRAO)

Is the chemical diversity at a 1000 au scale also present in the inner envelope/disk system (50 au) ?

What molecules are passed from the envelope to the disk in which planets, asteroids, and comets form

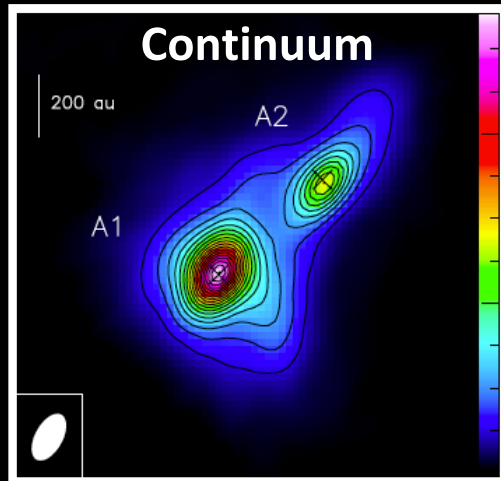
**STAY
TUNED !**
first papers
Bianchi+ 2020
De Simone+ 2020
Okoda+ 2021



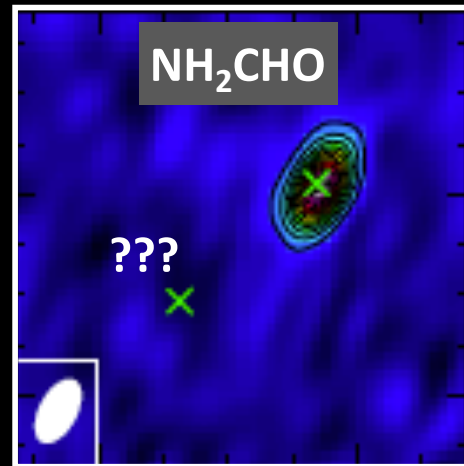
PPVII chapter



ACO: ALMA and beyond: Hot-corinos at cm-wavelengths



NGC1333-IRAS4A

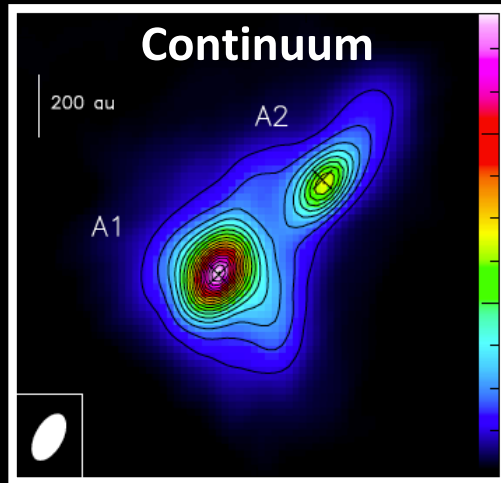


López-Sepulcre+ 2017;

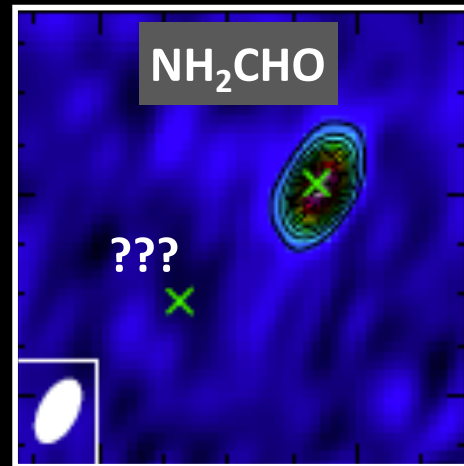
With ALMA:
Hot corino in one of the
two components

iCOMs abundances at
mm-wavelengths can be
underestimated

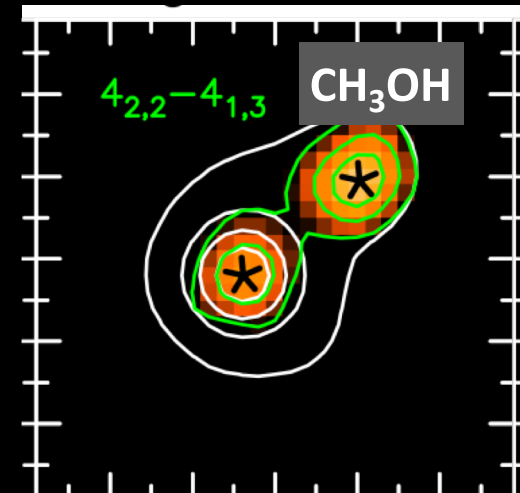
ACO: ALMA and beyond: Hot-corinos at cm-wavelengths



NGC1333-IRAS4A



López-Sepulcre+ 2017;



De Simone+ 2020

With ALMA:
Hot corino in one of the
two components

With VLA:
Both IRAS4A1 and IRAS 4A2
have a hot corino !

iCOMs abundances at
mm-wavelengths can be
underestimated

Food for SKA



ACO: ALMA and beyond: Hot-corinos at cm-wavelengths

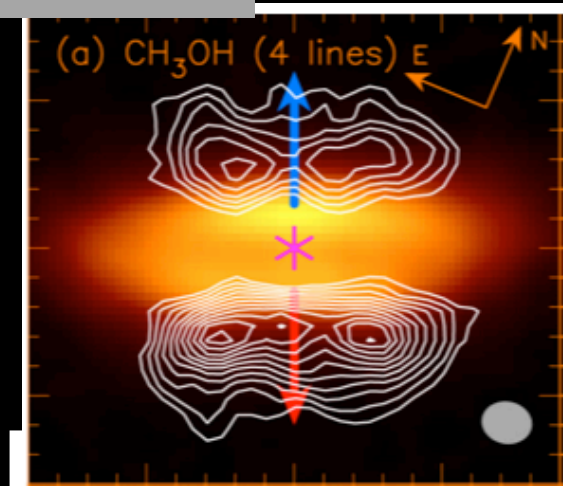
The disk midplane is optically thick in continuum and show
No iCOMs emission: lower abundance or opacity effect ?

Food for SKA
(WG Cradle of Life)

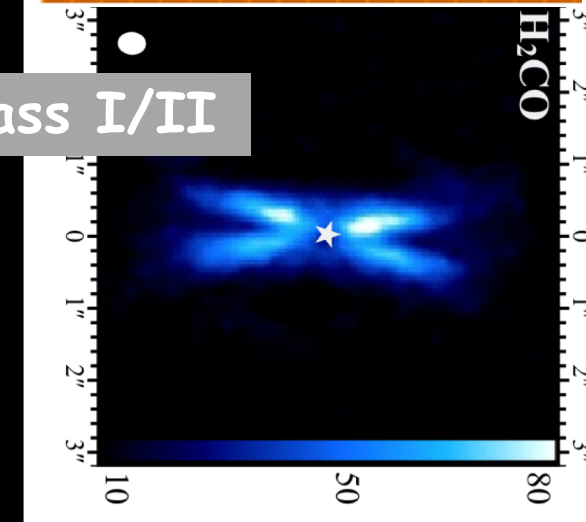


Lee+ 2019
Codella+ 2019

Class 0



Class I/II



Podio+ 2020



ACO for Ariel



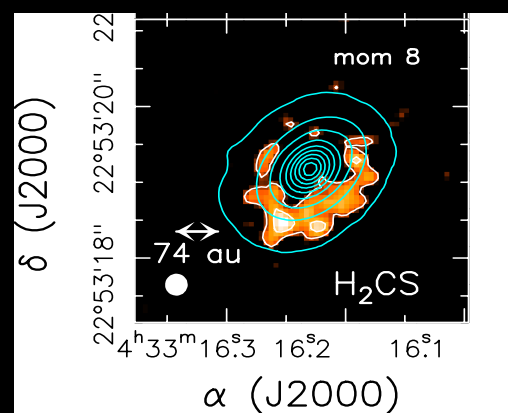
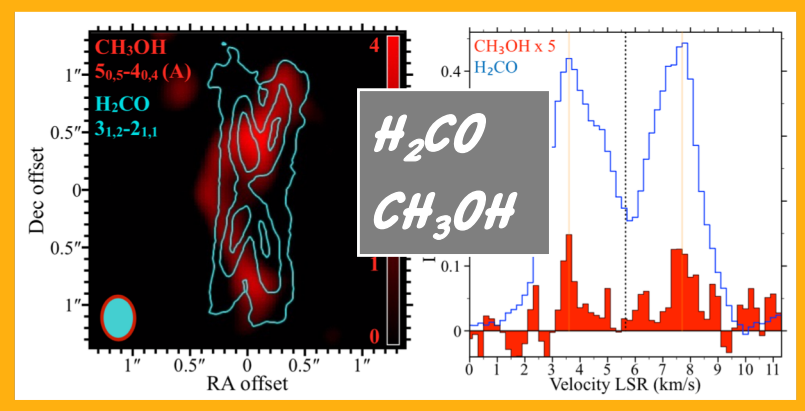
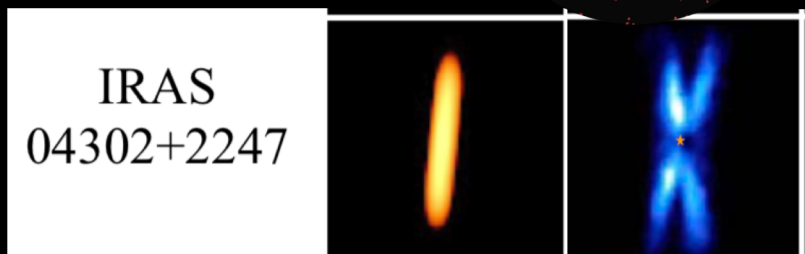
WP Planet Formation (Turrini+ 2021b - Ariel)

iCOMs in protoplanetary disks:
- To estimate their snowlines;

- To estimate the fraction of C, O, and N
trapped in organic molecules

- S/H ratio:

Class I SVS13A: 10^{-2} - 10^{-1}
($\times 10^{-5}$)



Podio+ 2020ab
Garufi+ 2020ab,
Codella+ 2020

Turrini+ 2021b

Object	S/H ($\times 10^{-5}$)
Sun	1.32
DM Tau	$< 10^{-2}$
GM Aur	
GO Tau	$< 10^{-2}$
HD 100546	$* 10^{-4}$ $* 10^{-3}$
IM Lup	
LkCa 15	$< 10^{-2}$
TW Hya	$* 10^{-4}$



Dissemination & Outreach

www.aco-itn.org

1° step - Visual Identity

2° step - <http://www.aco-itn.org/> – **here all the scientific hub.**

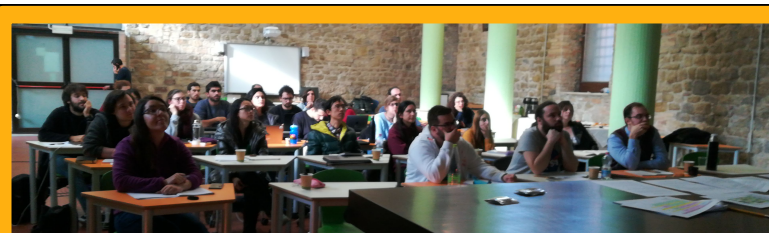
(outreach in English, French, Italian, Spanish, Catalan, Chinese... next step: Portuguese, German, and Arabic)

3° step – courses about communication for ESRs (**INAF + APRE**)

4° step – social FB and the **Chinese** Weibo

5° step – Torino [ACO DOC conference](#)

6° step – Virtual Reality school and Project Next July in Padova



Full PhDs engagement



THE SEEDS OF ACO

60 x 660



*I Congresso Nazionale di Astrochimica e
Astrobiologia (proto-)planetaria
21-23 October 2019 Duino (TS)*

ACO is just the start ! Plans for next 10 years and beyond

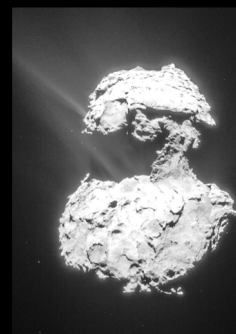


Outer Solar System Objects (OSSOs)

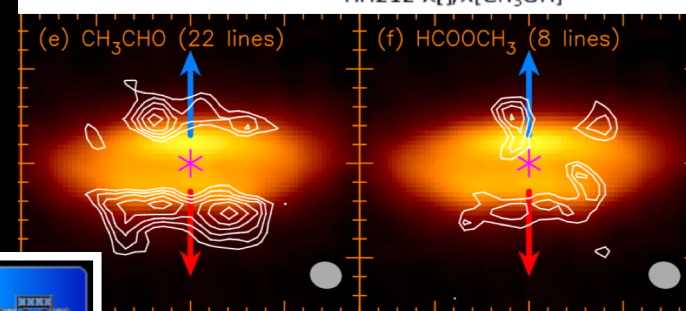
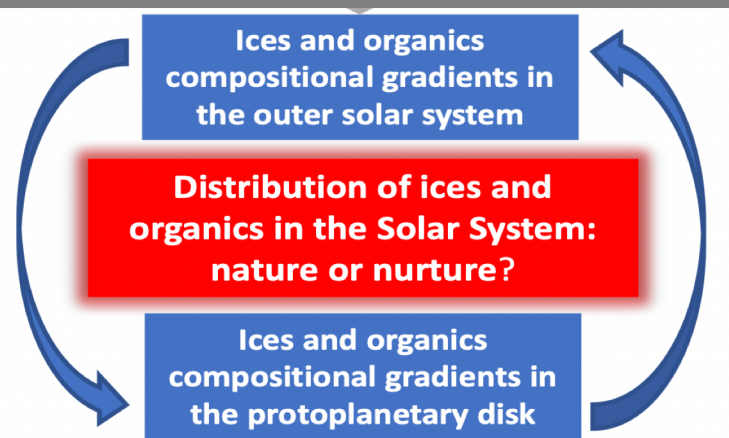
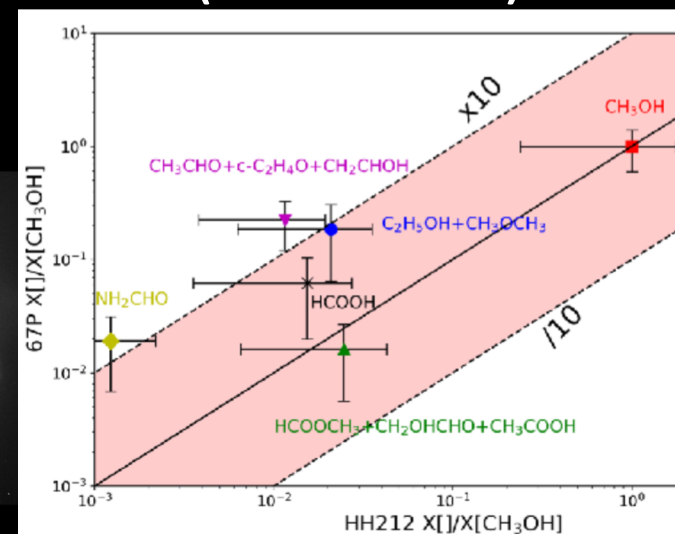


OSSOs: the most pristine material in the Solar System

Rosetta: The Italian community (INAF-IAPS in particular), provided a leading effort in VIS-IR reflectance spectroscopy experiments (VIRTIS-Rosetta, PI: F. Capaccioni)



Planetary composition:
inheritance ?
(Bianchi+ 2020)



Protostellar disk



ACO is just the start !

Plans for the next 10 years and beyond



AMBITION

ESA Voyage 2050 long-term
(2035-2050) plan

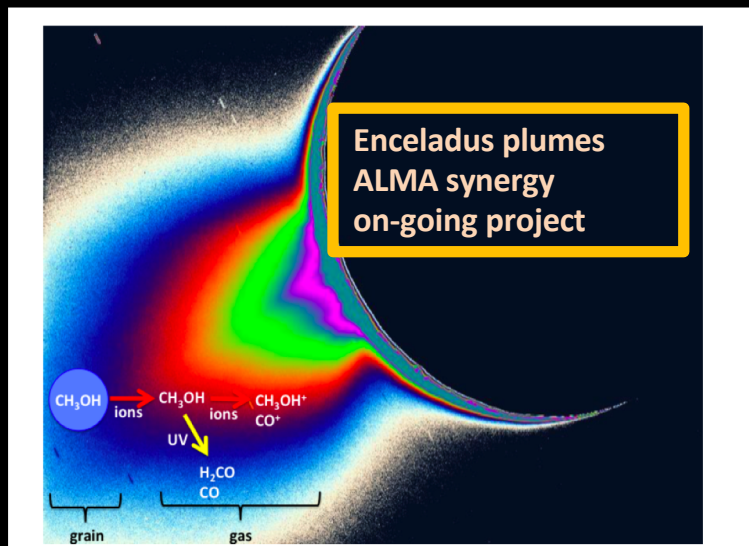
Bockelée-Morvan et al.
(2019 WB, 2021)



Cassini at the Saturn system,
Dawn at Vesta and Ceres
New Horizon at the Plutonian system and Kuiper Belt)
Juno (Jupiter and its moons)
Leading effort in VIS-IR reflectance spectroscopy
experiments onboard spacecrafts (VIMS-Cassini, VIR-
Dawn, PI: M.C. De Sanctis, JIRAM-Juno, PI: A. Mura)

This allows us to characterize the **compositional gradients**
in the outer Solar System;

Comparison with protosolar nebula by exploiting the
unique database of ALMA data probing the **disk chemical**
composition (FAUST, ALMA-DOT, VIPS)



ACO is just the start ! Plans for the next 10 years and beyond

Bridging exoplanet compositions and chemistry of protoplanetary disks

Exoplanet atmosphere

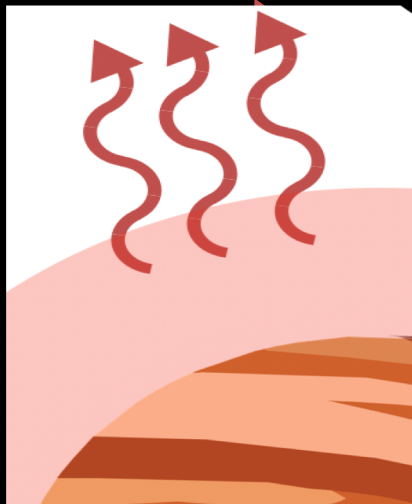
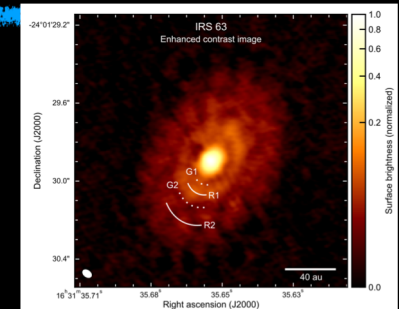
**Refractory elements
and molecules (Fe, TiO)**

**Volatile elements
and molecules (H₂O, CO)**

Protoplanetary disk

**Rocky component
of protoplanetary disk**

**Icy and gaseous component
of protoplanetary disk**

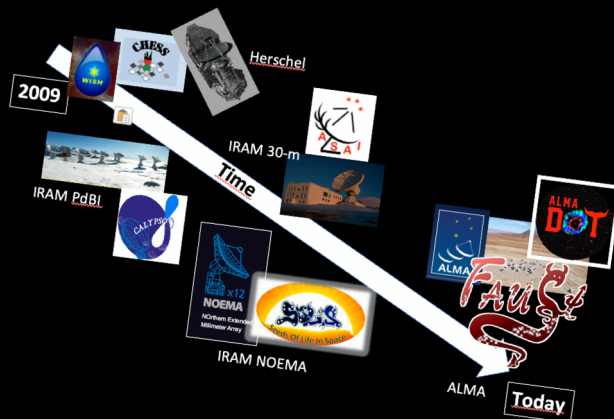


*We will perform the first
homogeneous
characterization of
refractory element
abundances in exoplanet
atmospheres, (Pino et al.
2020).*

*ongoing and future
surveys of volatile elements
(e.g. ARIEL, GAPS2).
HARPS-N@TNG
ESPRESSO@VLT
MAROON-X@Gemini-N
CARMENES@CAHA*

ACO is just the start ! Plans for the next 10 years and beyond

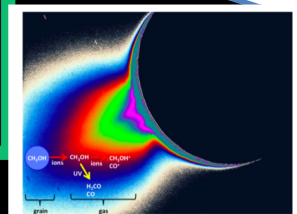
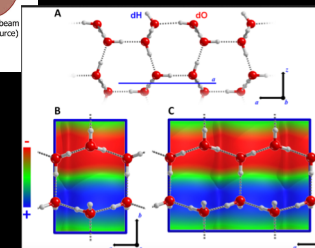
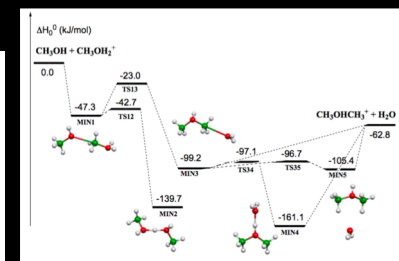
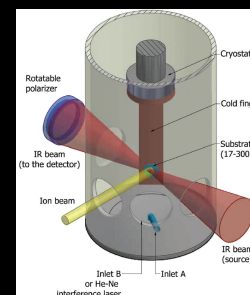
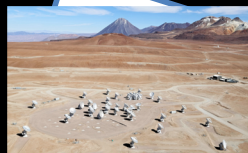
(i) laboratory outputs, (ii) surface simulations,
(iii) quantum-mechanical computations



VIPS ALMA large
program on disks
GBT to characterise
the C-chains in the
protostellar envelope

Chemical census of
protoplanetary disks with
VLA

ALMA characterisations of
Enceladus and Europa
plumes



...2030



Ariel, SKA, AMBITION

**INAF has a world-wide recognised leading role
in astrochemistry.**

A tight competition with several outstanding groups:

**CfA-Harvard-USA, CSIC-Madrid, NRAO-USA,
University of Copenhagen**

With the INAF's support, e.g.:

- **performing computers for data calibration,**
 - **maintaining high-level laboratories,**
 - **laboratory-observation shared PhDs,**
- **hiring young experts in astrochemistry,**
we can keep pushing!