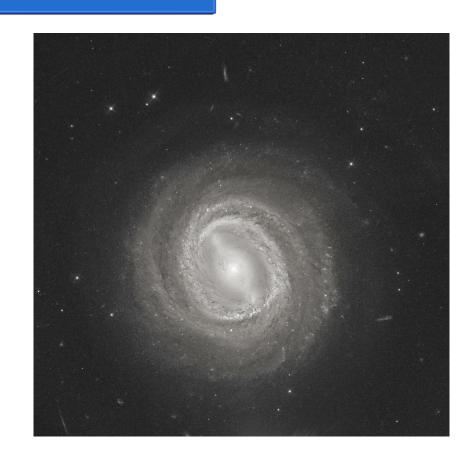
# LabX 2020-2021

What happened to the Seyfert 1.5 NGC 3783?



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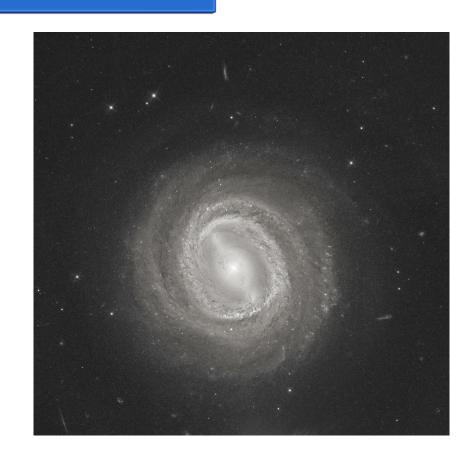
Target i.d.

Object type: G

Morphology: (R')SB(r)ab

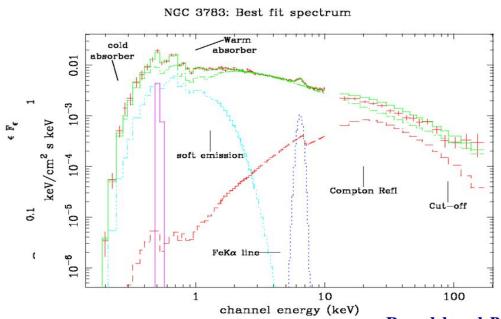
Activity Type: Sy 1.5

z≈0.009371



## What happened to the Seyfert 1.5 NGC 3783?

### The complex X-ray spectra

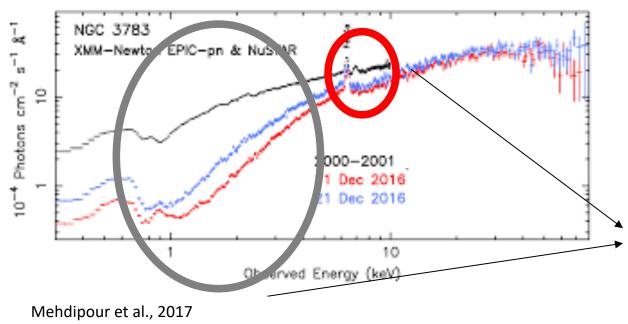


Very "typical" Seyfert 1 X-ray spectrum measured in all the X-ray observations

- 1) Very low cold absorption (in the UM scenario, it indicates that the putative dusty torus is not intercepting the line of sight. Ok for a type I AGN)
- 2) Fe line and reflection component detected → OK with the presence of accretion disk!
- 3) High-E cutoff measured at E≈100 keV: OK with thermal Comptonization
- 4) warm absorber measured: → ok with UM and presence of warm electrons along the polar axis of the system to explain measurement of broad lines in polarized light (optical)

Broad-band *BeppoSAX* observation NGC 3783. De Rosa et. al 2002

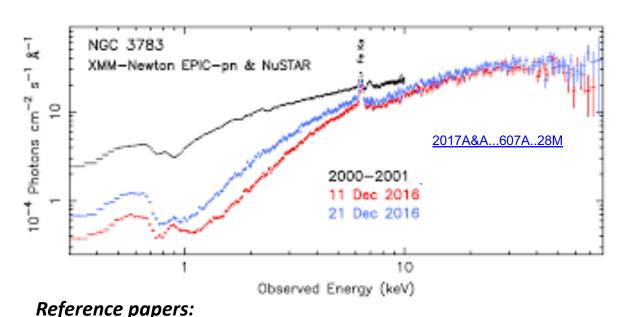
# What happened to the Seyfert 1.5 NGC 3783?



#### Goals

- 1) What are the properties of the primary emission of NGC 3783?
- 2) What are the properties of the absorbers in NGC 3783?
- 3) Are these properties in agreement with the predictions of UM for AGN?
  - ... but something changed... (optional part)
- 4) the changes were driven by the primary emission and/or the absorber?
- 5) can I interpret these changes within the UM scenario? How?

# What happened to the Seyfert 1.5 NGC 3783?



De Rosa et al., 2002, A&A, 387, 838 Mehdipour et al., 2017, A&A, 607, 28

#### How

#### **Mandatory part**

- a) use only one XMM-Newton observation
- b) use only EPIC/pn data in the 3-10 keV band
- c) perform data reduction and spectral analysis to infer:
- c-1) shape of the primary continuum;
- c-2) column densities and ionization states of the absorbers;
- c-3) dimensions of the regions where the Fe line is produced;
- c-4) reflection?
- c-5) something else?

#### **Optional part:**

d) re-do everything on the other observation, then compare and discuss the differences!